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# Numerical modelling of the quantum-tail effect on fusion rates at low energy<sup>☆</sup>

M. Coraddu<sup>a,b</sup>, G. Mezzorani<sup>a,b</sup>, Yu.V. Petrushevich<sup>c</sup>,  
P. Quarati<sup>d,b</sup>, A.N. Starostin<sup>c,\*</sup>

<sup>a</sup>Physics Department, University of Cagliari, Monserrato I-09042, Italy

<sup>b</sup>I.N.F.N. Cagliari, Monserrato I-09042, Italy

<sup>c</sup>State Research Center of Russian Federation, Troitsk Institute for Innovation and Fusion Research,  
Center for Theoretical Physics and Computational Mathematics, Troitsk, Moscow region 142190, Russia

<sup>d</sup>Physics Department, Politecnico Torino, Torino I-10125, Italy

## Abstract

Results of numerical simulations of fusion rate  $d(d,p)t$ , for low-energy deuteron beam, colliding with deuterated metallic matrix (Phys. Lett. B 547 (2002) 193; Eur. Phys. J. A 13 (2002) 377) confirm analytical estimates given in M. Coraddu et al. (Quantum-tail effect in low energy d+d reaction in deuterated metals, Physica A, this issue), taking into account quantum tails in the momentum distribution function of target particles, and predict an enhanced astrophysical factor in the 1 keV region in qualitative agreement with experiments.

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## 1. Introduction

Significant divergence from theoretical predictions of non-resonant fusion cross-section at low energies of incident deuteron particles has recently been observed in

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\* Corresponding author. Troitsk Institute for Innovation and Fusion Research (TRINITI), Troitsk, Moscow region 142092, Russian Federation. Fax: +7-095-334-5128.

E-mail address: [a.starostin@relcom.ru](mailto:a.starostin@relcom.ru) (A.N. Starostin).

1 experimental works [1,2]. At energies of the charged deuteron beam less than 5 keV,  
 2 colliding with deuterated metallic matrix, a great enhancement of the fusion cross-  
 3 section takes place, compared to theoretical evaluations. The hypothesis of ions inter-  
 4 acting via screened potential inside the metallic target was proposed in Ref. [1], but an  
 5 unrealistically large screening potential was necessary to obtain reasonable agreement  
 6 with the observed data. It is obvious that problems arising from experimental results  
 7 need refining the theoretical models actually used.

8 As it was shown in papers [3,4], the role of quantum corrections to the particle mo-  
 9 mentum distribution is quite important. A significant deviation from Maxwellian distri-  
 10 bution function appears. At large values of momentum, due to quantum corrections,  
 11 the tail of the distribution has a power-law asymptotic behavior instead of exponential,  
 12 resulting in modified reaction rates. In particular, this effect leads to non-exponential  
 13 temperature dependence of inelastic process rates at relatively low temperatures and  
 14 high densities.

15 In this paper we present results of numerical calculations of reaction rates for con-  
 16 ditions of the interacting particles in the beam and in the target such as those of the  
 17 experiments [1,2].

## 2. Numerical modeling of reaction rates for experimental conditions

18 The state of the system defined by the generalized distribution function over energy  
 19 and momentum  $F(E, \vec{p})$ , may be presented in the factorized form:  $F(E, \varepsilon) = n(E)a(E - \varepsilon)$ ,  
 20 where  $\varepsilon = p^2/2m$  is the particle kinetic energy. The reaction-rate constants of the  
 21 inelastic process between two particles, named “a” and “b”, may be presented in a  
 22 more general form by the integral (see Ref. [5], Starostin et al., this issue)

$$\begin{aligned}
 N_a N_b K_{ab} = & 8\pi C \int_{-\infty}^{\infty} dE_a \int d\vec{p}_a \int_{-\infty}^{\infty} dE_b \int d\vec{p} \int d\vec{q} \\
 & \times n(E_a)(1 - n(E_a + Q_a - \omega))a_a(E_a - \varepsilon_a) \\
 & \times n(E_b)(1 - n(E_b + \omega + Q_b))a_b(E_b - \varepsilon_b) \\
 & \times a'_a(E_a + Q_a - \omega - \varepsilon_{\vec{p}_a - \vec{q}})a'_b(E_b + \omega + Q_b - \varepsilon_{\vec{p}_b + \vec{q}})\sigma(\varepsilon_p)\sqrt{\frac{2\varepsilon_p}{M}}, \quad (1)
 \end{aligned}$$

23 where  $E_a$ ,  $p_a$  are the energy and momentum of the “a” particle,  $\varepsilon_p = p^2/2M$  is the  
 24 relative kinetic energy in the center of mass,  $M$  is the reduced mass of colliding  
 25 particles,  $C$  is a normalization constant, defined from comparison of the expression  
 26 calculated by (1) and known results at high temperature and low density. We have

$$\vec{p} = \frac{m_b \vec{p}_a - m_a \vec{p}_b}{m_a + m_b}.$$

27 The expression for the population number  $n(E)$  depends on the statistical distribution  
 28 of the system. For the purpose of this work we must consider that deuterons, which  
 29 are bosons, have the distribution

$$n(E) = (\exp[(E - \mu)/T] - 1)^{-1},$$

1 where  $\mu$  is chemical potential. For non-ideal plasma, the spectral dependence of the  
 2 distribution function, defined by Lorentzian profile, is

$$a(E, \vec{p}) = \frac{\gamma_i(E, \vec{p})}{\pi[(E - \varepsilon_p - \Delta(E, \vec{p}))^2 + \gamma_i^2(E, \vec{p})]} \quad (2)$$

3 In Eq. (2)  $\gamma$  is the width,  $\Delta$  stands for the energy shift due to atom–matter interaction.  
 4 The line width is given by  $\gamma_a = N\hbar\sigma_a V_a$ ; where  $\sigma_a = \pi e^4/\varepsilon_a^2$ ,  $V_a = \sqrt{2\varepsilon_a/m_a}$ ,  $N$  is the  
 5 concentration of scattering centers. For ideal plasma conditions, i.e., when the density  
 6 decreases, the width  $\gamma(E, \varepsilon_p) \rightarrow 0$ , the function  $a(E, \vec{p})$  becomes a delta-function. The  
 7 cross-section dependence on kinetic energy may be given in the form

$$\sigma_0(\varepsilon_p) = \frac{S(\varepsilon_p)}{\varepsilon_p} \exp(-2\pi\eta(\varepsilon_p)), \quad \eta(\varepsilon_p) = \frac{Z_1 Z_2 e^2}{\hbar} \sqrt{\frac{M}{2\varepsilon_p}} \quad (3)$$

8 where  $\eta(\varepsilon_p)$  is the Sommerfeld parameter. The astrophysical factor  $S$  is weakly  
 9 dependent on the kinetic energy.

10 The influence of the screening potential  $U_e$  on the reaction rate, due to the effect of  
 11 metallic electrons, may be taken into account by adding  $U_e$  to the collision energy

$$\sigma(\varepsilon_p) = \sigma_0(\varepsilon_p + U_e) \quad (4)$$

12 Numerical calculations of the reaction rates were performed in accordance with the  
 13 above model, for conditions close to the experimental ones: target particles concentra-  
 14 tion  $N_a = 5 \times 10^{23} \text{ cm}^{-3}$ , and interacting particle masses:  $m_a = m_b = 2 \text{ amu}$ . Fusion  
 15 reactions were considered between target particles of kind “a” and beam particles of  
 16 kind “b”. In the expression of the width we used as concentration  $N$  the concentration  
 17 of the scattering ions in the metallic matrix.

18 Taking into account high dimension, the computation of the integral (1) can be  
 19 performed using Monte Carlo method. The kinetic energy distribution of target particles  
 20 was taken at temperature  $T = 2.44 \times 10^{-2} \text{ eV}$ , while the “b” particles were taken with  
 21 the beam energies. It is interesting to note that for the case of a beam of mono-energetic  
 22 particles the reaction rate computation can be reduced from expression (1) to the more  
 23 simple integral

$$N_a K' = C \int_0^\infty dE_a \int d\vec{p} n_a(E_a) a(E_a - \varepsilon_p, \varepsilon_p) \sqrt{\frac{2\varepsilon_p}{M}} \sigma(\varepsilon_p) d\varepsilon_p \quad (5)$$

24 Within the framework of such model, for the case of ideal plasma, when we can  
 25 neglect the wings of the Lorentzian profile in the integral (5), the expression for  
 the reaction rate can be further simplified:

$$\begin{aligned} N_a K_2 = N_a K' &= C \int_0^\infty dE_a \int d\vec{p} \int n_a(E_a) \delta(E_a - \varepsilon_p) \sqrt{\frac{2\varepsilon_p}{M}} \sigma(\varepsilon_p) d\varepsilon_p \\ &\sim \int_0^\infty n_a(\varepsilon_p) \varepsilon_p \sigma(\varepsilon_p) d\varepsilon_p \end{aligned} \quad (6)$$

26 The influence of the distribution wings on the reaction rate value can be obtained by  
 27 comparison of the computation results of the two expressions (1) and (5). We can

1 also compare such results with the calculated reaction rate  $K_1 = \sigma V$ , using expression  
 2 (3). After such comparison it is possible to estimate the astrophysical factor  $S$  and the  
 3 deviation of theoretical predictions from experimental data.

4 To estimate the influence of momentum distribution tails on the reaction rate and  
 5 the difference with the Maxwellian case, it is necessary to take into account the finite  
 6 width of the Lorentzian profile (2). As was shown in Refs. [4,5], the main result of  
 7 quantum corrections is that the momentum distribution function has asymptotically a  
 power-law tail

$$f(\varepsilon_p) = C' \int_0^\infty dE_a a(E_a - \varepsilon_p, \varepsilon_p) \sim \exp(-\varepsilon_p/T) + C_a(T)/\varepsilon_p^4. \quad (7)$$

9 Using such decomposition in Eq. (5), it is possible to calculate the reaction rate, taking  
 into account non-Maxwellian distribution function

$$K_3 = C_3 \int_0^\infty d\varepsilon_a f(\varepsilon_a) \sqrt{\frac{\varepsilon_p \varepsilon_a}{M}} \sigma(\varepsilon_p). \quad (8)$$

11 The results of such calculations are shown in the Table 1.

12 The reaction rate constants, calculated by different models, agree among themselves  
 13 at beam energies above 2 keV. Decreasing energy in the range between 2 and 1 keV,  
 14 we have found that  $K_1$  and  $K_2$  have still close values, while the rate constant  $K$  is much  
 15 larger. It is interesting to note that the constants  $K_3$  and  $K$  have relatively close values.  
 16 Thus, it is possible to conclude, that, for correct estimations of the rate constants,  
 17 it is quite possible to use the expressions shown in Eq. (5). The results of these  
 18 calculations show that the wings of the momentum distribution are very important for  
 19 a correct evaluations of the reaction rates. In the last column of Table 1 we presented  
 the factor which characterizes the deviation of the rate in cause of non-ideal plasmas.

21 It is interesting also to estimate the role of the screening effect on the reaction rate  
 and to compare it to the considered mechanisms. For that purpose the calculations  
 23 were performed within the framework of the proposed model, but with addition of  
 a screening potential  $U_e = 28$  eV. Its influence was taken into account in accordance  
 25 with expression (4). Such value of the potential seems realistic for the experimental  
 conditions [1]. The results of calculations are presented in the Table 2.

Table 1

Comparison of the reaction rates  $\langle \sigma v \rangle$  as function of the energy of the beam using the general expression  
 of Eq. (1),  $K$ , and the three models in Eq. (3),  $K_1$ , in Eq. (5),  $K_2$ , and in Eq. (8),  $K_3$

$E_a$ (keV)	$K_1$	$K_2$	$K_3$	$K$	$K/K_1$
15	4.381E+04	4.045E+04	7.393E+04	4.38E+04	1.00E+00
10	4.073E+03	3.762E+03	6.877E+03	4.11E+03	1.01E+00
5	1.711E+01	1.580E+01	2.892E+01	1.77E+01	1.03E+00
2	2.615E - 04	2.421E - 04	4.487E - 04	2.85E - 04	1.09E+00
1.8	5.038E - 05	7.223E - 05	1.344E - 04	5.62E - 05	1.12E+00
1.5	2.339E - 06	3.850E - 06	7.343E - 06	3.34E - 06	1.43E+00
1.2	3.613E - 08	7.474E - 08	2.265E - 07	7.84E - 07	2.17E+01
1	8.252E - 10	7.711E - 10	5.678E - 08	2.82E - 07	3.42E+02

Table 2

Same as Table 1 taking into account the screening effect according to Eq. (4) with  $U_e = 28$  eV

$E_a$ (keV)	$K_1$	$K_2$	$K_3$	$K_3/K_1$
15	4.474E+04	4.133E+04	7.552E+04	1.6879937
10	4.237E+03	3.914E+03	7.152E+03	1.6880552
5	1.911E+01	1.766E+01	3.232E+01	1.69143
2	4.059E-04	5.370E-04	9.931E-04	2.4466201
1.8	8.433E-05	1.184E-04	2.197E-04	2.6056397
1.5	4.605E-06	7.336E-06	1.383E-05	3.0034846
1.2	9.315E-08	1.819E-07	4.320E-07	4.638S3085
1	2.863E-09	6.951E-09	7.386E-08	25.801192

1 Here we have not presented the results of the more general computations using  
 2 Eq. (1), because we have obtained rather correct estimates using model (5). We find  
 3 a weak influence of the screening effect, using the reasonable value of the potential,  
 4 which agrees with the results of [1,2]. Taking into account quantum corrections, the  
 5 theoretical evaluations of the rates increase, in the low-energy range at 1–2 keV and  
 less, if compared to the rates evaluated without the quantum effect.

### 7 3. Uncited reference

[6].

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